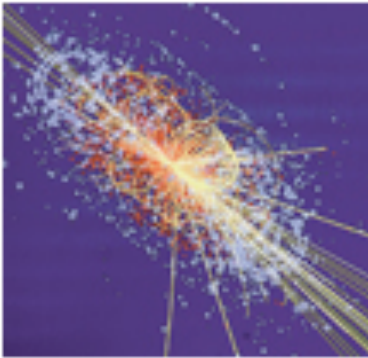
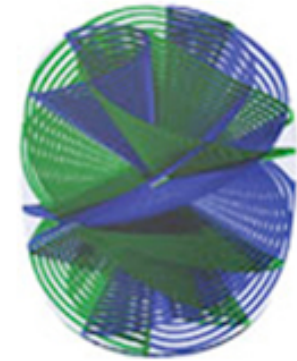

Indirect Constraints and SUSY



*Albert De Roeck
CERN &
Antwerp University &
IPPP Durham, UK*

*Oliver Buchmuller
CERN*



LHC New Physics Signatures Workshop

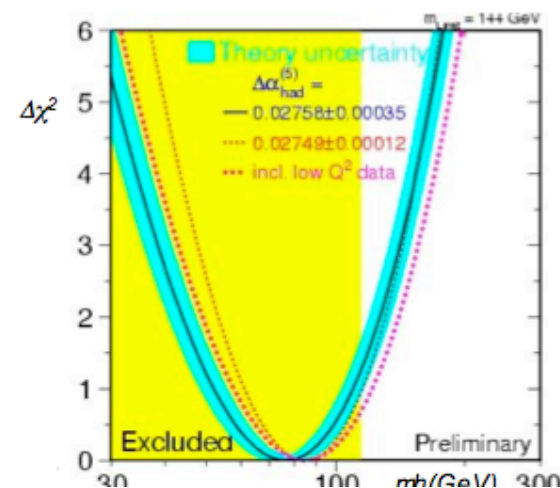
Use of indirect constraints: EW Fit

A prominent example for the comprehensive usage of indirect constraints!



Use high precision EW measurements to:

- Test the consistency of the SM
- Predict the last unknown quantity in the SM: the higgs mass



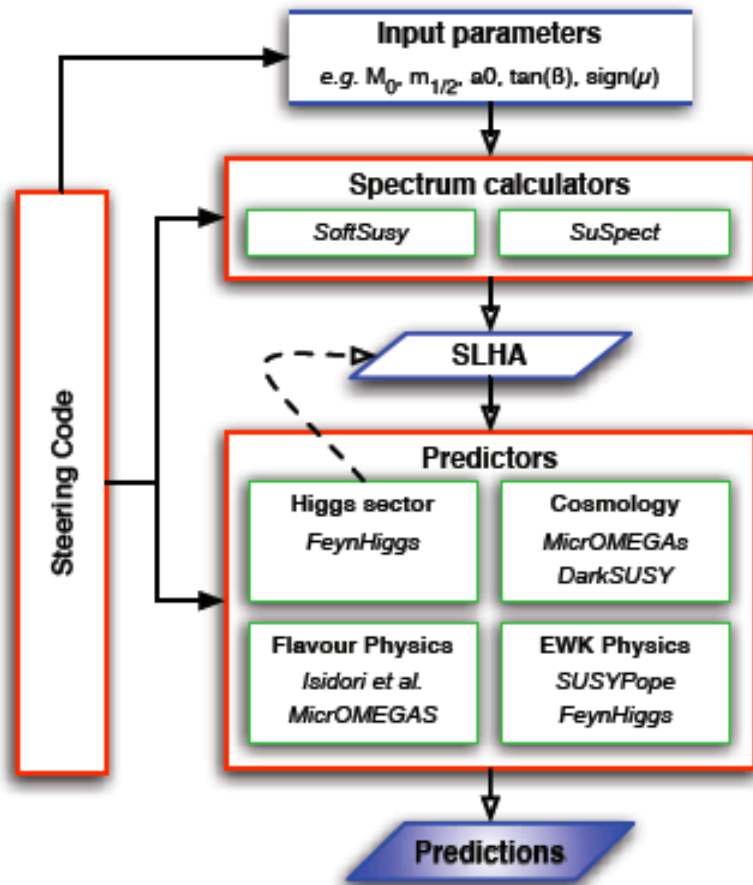
Common framework



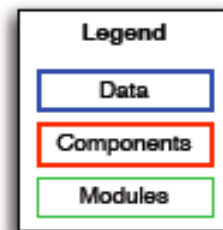
- Goal: a framework to provide consistent indirect constraints
- Collaboration of interested theorists and experimentalists
 - Buchmüller, Oliver (CERN) – Exp.
 - De Roeck, Albert (CERN & Uni. Antwerpen) – Exp.
 - Heinemeyer, Sven (Santander) – Theo.
 - Olive, Keith (Uni. of Minnesota) – Theo.
 - Ronga, Frédéric (CERN) – Exp.
 - Weiglein, Georg (Durham) – Theo.
 - Cavanaugh, Richard (Uni. of Florida) – Exp.
 - Ellis, John (CERN) – Theo.
 - Isidori, Gino (INFN Frascati) – Theo.
 - Paradisi, Paride (Uni. of Valencia) – Theo.
 - Weber, Arne (Max Planck Inst. (Munich)) – Theo.
- Started at workshop on [Flavour Physics in the Era of the LHC](#)
 - ⇒ See (draft) report, sec. 5.2
- Main focus of the work:
 - Development of a *common tool* for indirect constraints
 - Compilation (and integration) of state-of-the-art predictions
 - Application of the tool

Buchmuller et al., PLB 657/1-3 pp 87-94

Common framework



- Consistency
Relies on SLHA interface
 - Modularity
Compare calculations
Add/remove predictions
 - State-of-the art calculations
Direct use of code from experts
- ⇒ A unique "platform" for the integration of tools



First Exercise:
Look for regions in
CMSSM Space
using the constraints

“Preferred” Parameter Space

Pulls from mSUGRA fit:
 $\chi^2/\text{NDF} = 17/14$; $P(\chi^2)=20\%$

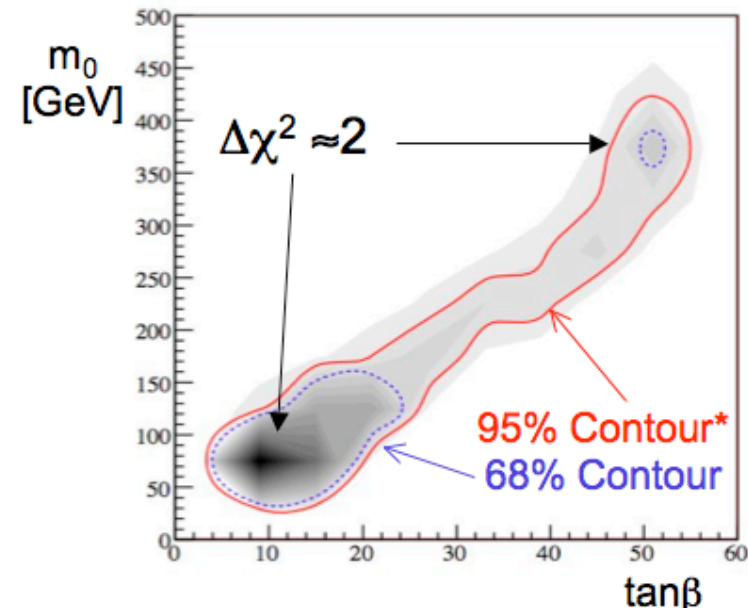
Variable	Measurement	Fit	$ O^{\text{meas}} - O^{\text{fit}} /\sigma^{\text{meas}}$
$\Delta\alpha_{\text{had}}^{(5)}(m_Z)$	0.02758 ± 0.00035	0.02774	0.1
m_Z [GeV]	91.1875 ± 0.0021	91.1873	0.1
Γ_Z [GeV]	2.4952 ± 0.0023	2.4952	0.1
σ_{had}^0 [nb]	41.540 ± 0.037	41.486	1.4
R_1	20.767 ± 0.025	20.744	0.9
$A_{\text{fb}}^{0,l}$	0.01714 ± 0.00095	0.01641	0.8
$A_1(P_\tau)$	0.1465 ± 0.0032	0.1479	0.1
R_b	0.21629 ± 0.00066	0.21613	0.1
R_c	0.1721 ± 0.0030	0.1722	0.1
$A_{\text{fb}}^{0,b}$	0.0992 ± 0.0016	0.1037	2.8
$A_{\text{fb}}^{0,c}$	0.0707 ± 0.0035	0.0741	1.1
A_b	0.923 ± 0.020	0.935	0.5
A_c	0.670 ± 0.027	0.668	0.1
$A_1(\text{SLD})$	0.1513 ± 0.0021	0.1479	1.5
$\sin^2\theta_{\text{eff}}^{\text{lept}}(Q_{\text{fb}})$	0.2324 ± 0.0012	0.2314	0.8
m_W [GeV]	80.398 ± 0.025	80.382	0.6
m_t [GeV]	170.9 ± 1.8	170.8	0.1
$R(b \rightarrow s\gamma)$	1.13 ± 0.12	1.12	0.1
$B_s \rightarrow \mu\mu$ [$\times 10^{-8}$]	< 8.00	0.33	N/A (upper limit)
Δa_μ [$\times 10^{-9}$]	2.95 ± 0.87	2.95	0.1
Ωh^2	0.113 ± 0.009	0.113	0.1

Collaboration of experimentalist and theorist:
arXiv:0707.3447

BUCHMULLER, CAVANAUGH, DE ROECK, HEINEMEYER,
 ISIDORI, PARADISI, RONGA, WEBER, WEIGLEIN.

$$\chi^2 = \sum_i^N \frac{(C_i - P_i)^2}{\sigma(C_i)^2 + \sigma(P_i)^2} + \sum_i^M \frac{(f_{\text{SM}_i}^{\text{obs}} - f_{\text{SM}_i}^{\text{fit}})^2}{\sigma(f_{\text{SM}_i})^2}$$

Multi-parameter fit using all mSUGRA parameters. Relevant SM uncertainties like Δm_{top} are also considered

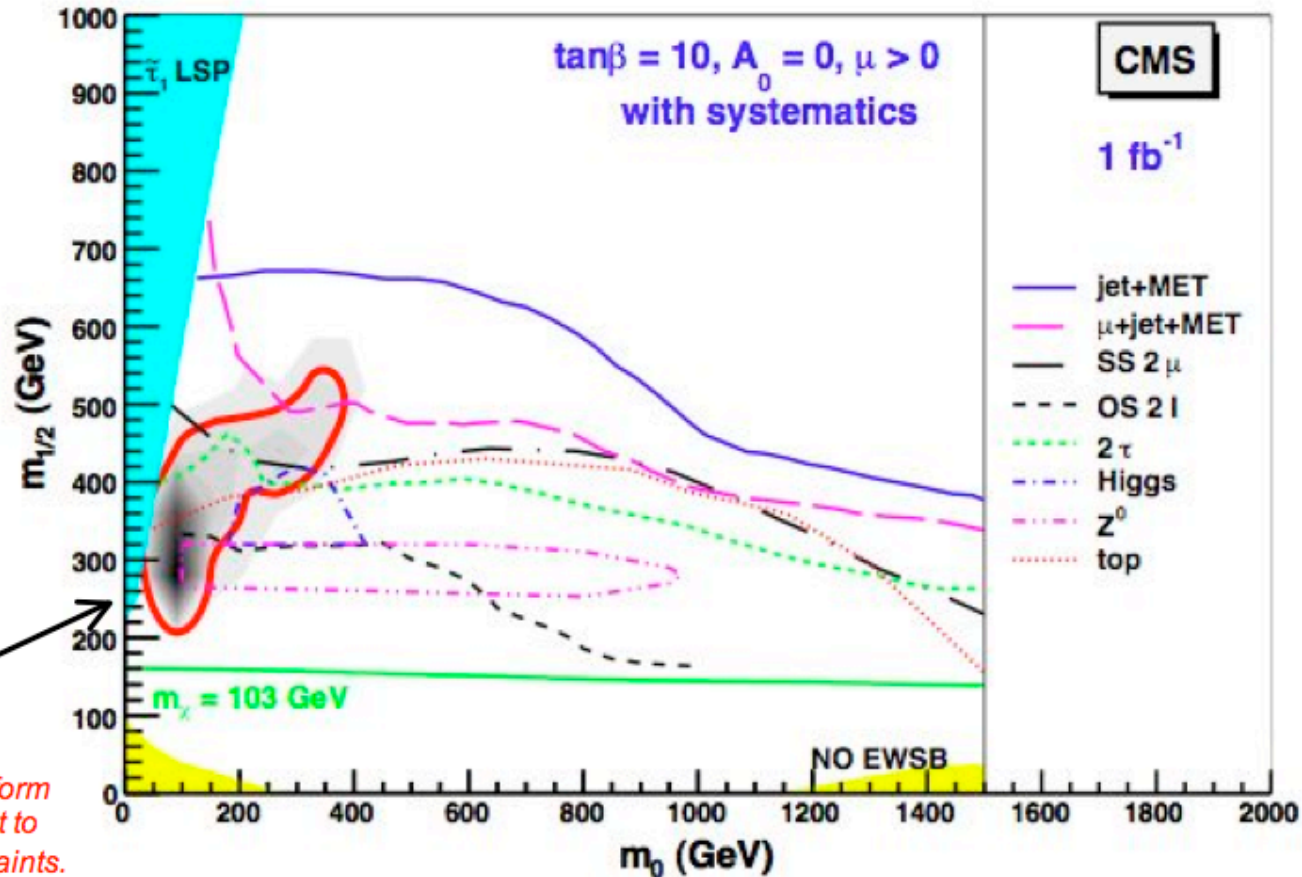


Preferred SUSY Parameter Space

CMS Discovery Reach for 1fb^{-1} (ATLAS similar)

Example of similar analyses:

- Ellis, Heinemeyer, Olive, Weber, Weiglein - ph/0706.0652
- Allanach, Lester, Weber - ph/0705.0487
- Trotta, Austri, Roszkowski - ph/0609126
- ... there are more!



arXiv:0707.3447

*95% contour obtained from a multi-parameter χ^2 fit to important indirect constraints.
 $\chi^2/\text{NDF} = 17/15$ - good fit*

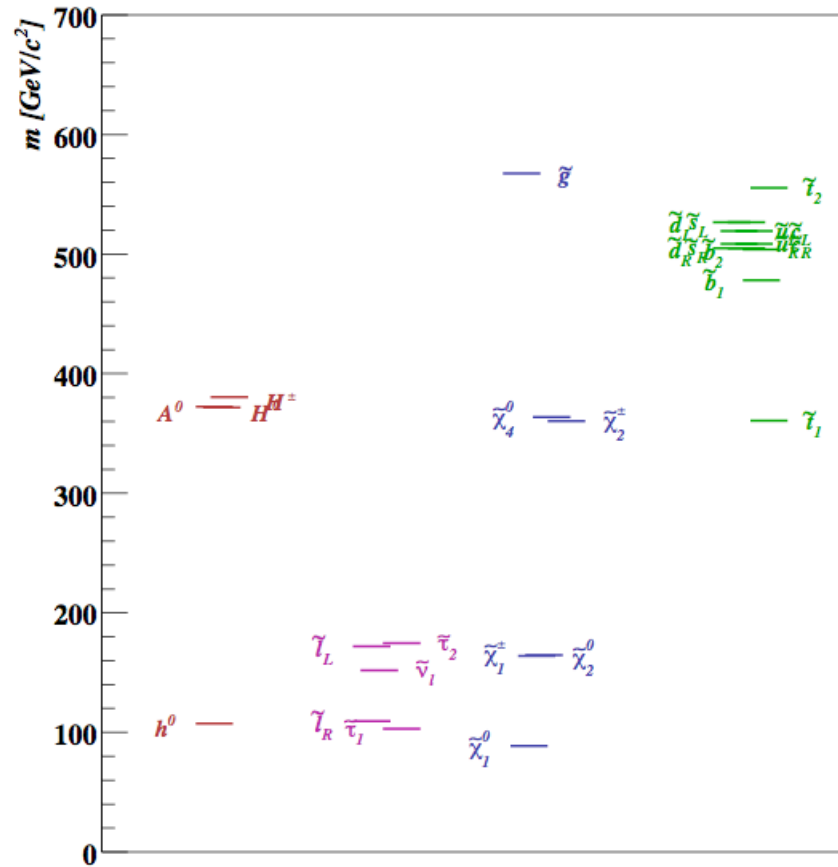
NOTE: All $m\text{SUGRA}$ parameters are free in the fit!

“CMSSM fit clearly favors low-mass SUSY”

Sparticle Spectrum



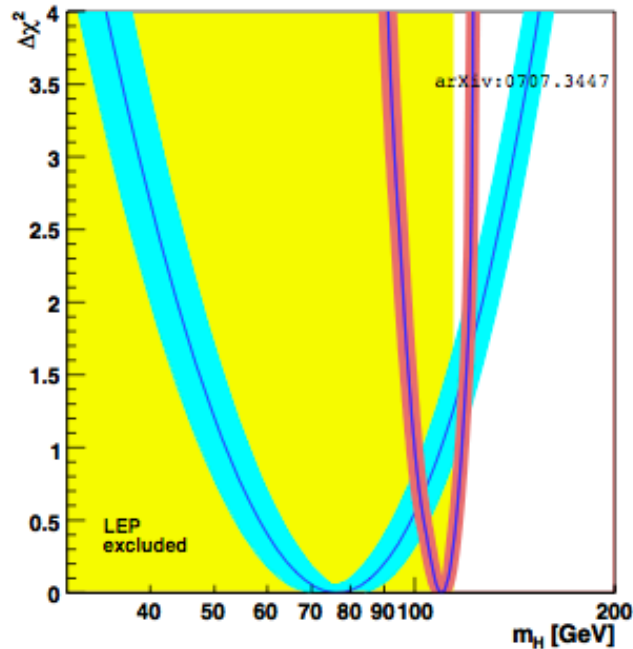
“best CMSSM Fit”



“LHC weather
forecast”

M0 M12 A0 tb
 49.2 232.3 -122.4 6.9
 Ma=372 GeV; mu=336 GeV; mh=111 GeV

Higgs Mass “Prediction”



LEP m_H constrained not used in fit

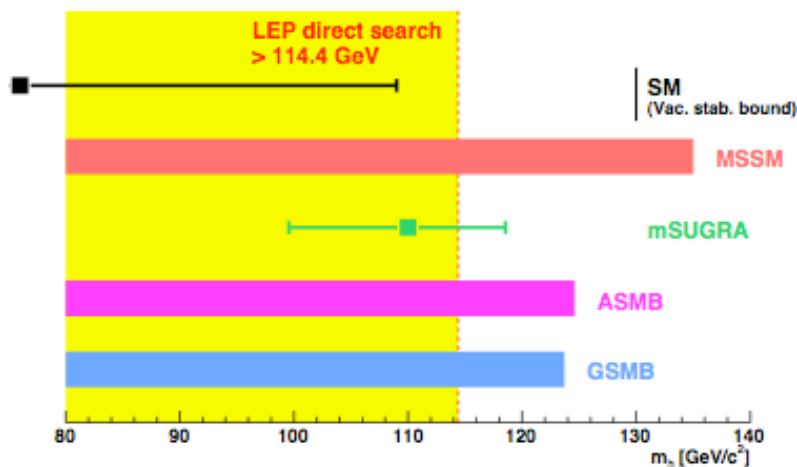
- Constrain m_h to scan value;
 - minimize all model parameters in each point;
- ⇒ determine error on m_h prediction

SM fit:

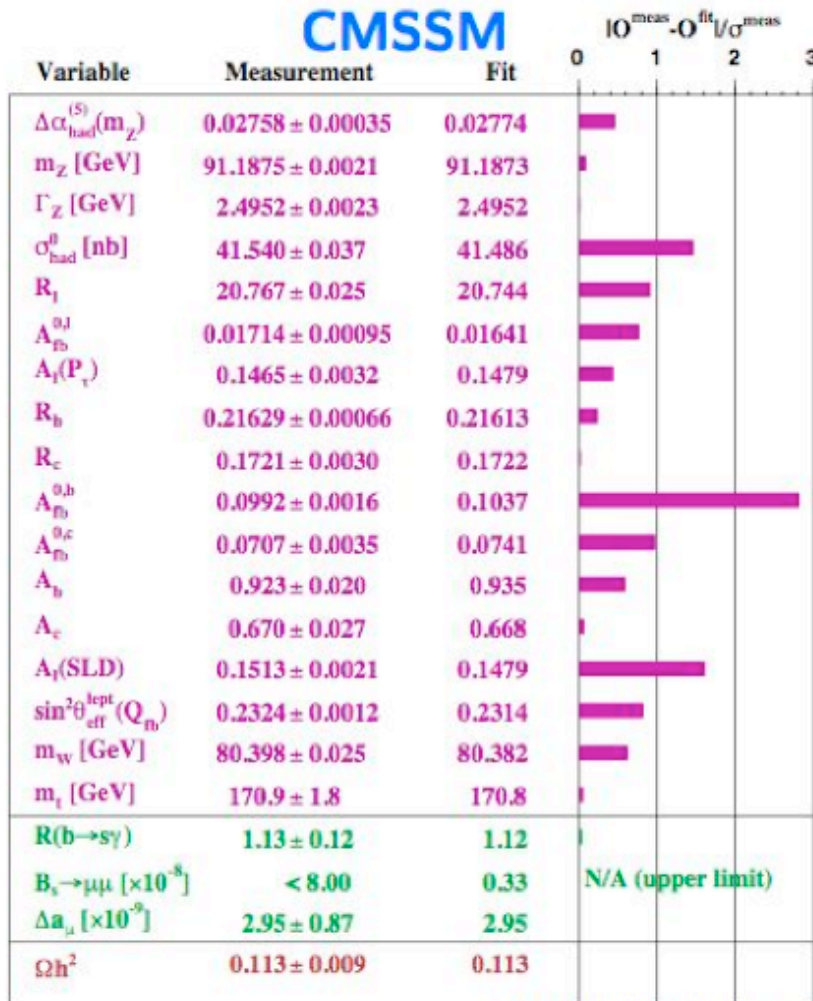
- $m_H = 78^{+33}_{-24} \text{ GeV}/c^2$
- 12% probability at exclusion limit
Including theoretical uncertainty

CMSSM fit:

- $m_h = 110^{+8}_{-10} \pm 3 \text{ GeV}/c^2$
- 20% probability at exclusion limit
Including theoretical uncertainty

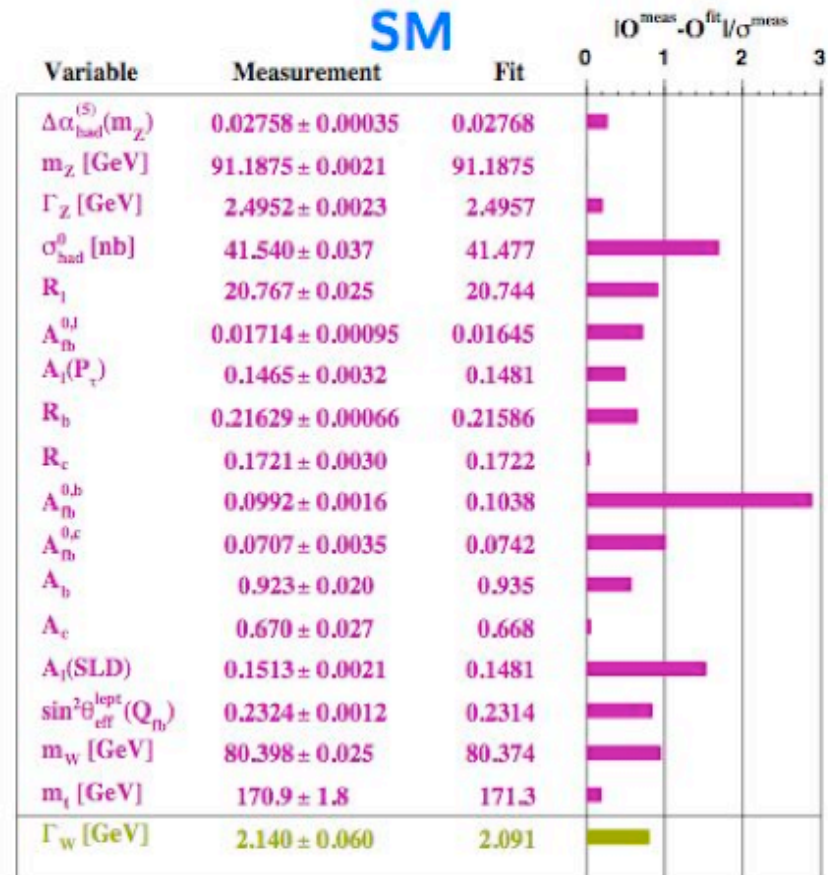


CMSSM vs. SM



arXiv:0707.3447 [hep-ph]

$\chi^2/\text{ndof} = 17.0/13$ (20% prob.)



arXiv:hep-ex/0612034

$\chi^2/\text{ndof} = 18.2/13$ (15% prob.)

New Analyses



Markov Chain Monte Carlo (MCMC):

To ensure a comprehensive mapping of the parameter space we have performed several MCMC's with many different starting points. The shown contour is the combined result of all of them.

χ^2 Minima:

The overall χ^2 minima is determined using Minuit. The chosen starting values are determined from the results of the MCMC sampling of the parameter space.

Contours:

Contours are defined from all MCMC's. So far, we have not performed toys to validate and refine all the 68% (blue) and 95% (red) contours but cross checks show that the contours are reliable.

Extension of the Constraints



Low energy observables

$BR(b \rightarrow s\gamma)$	MicrOMEGAs	Isidori & Paradisi
$BR(b \rightarrow sll)$		Isidori & Paradisi
$BR(B_s \rightarrow \mu\mu)$	MicrOMEGAs	Isidori & Paradisi
$BR(B \rightarrow \tau\nu)$		Isidori & Paradisi
$BR(K \rightarrow \tau\nu)$		Isidori & Paradisi
$BR(K \rightarrow \pi\nu\nu)$		Isidori & Paradisi
$\Delta m_s / \Delta m_d$		Isidori & Paradisi
Δm_s		Isidori & Paradisi
Δm_K		Isidori & Paradisi
$g-2$	FeynHiggs	

Higgs sector observables

m_h^{light}	FeynHiggs
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Cosmology observable

Ωh^2	MicrOMEGAs	DarkSUSY
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High energy EW observables


R_l	A. Weber <i>et al.</i>
R_b	A. Weber <i>et al.</i>
R_c	A. Weber <i>et al.</i>
$A_{fb}(b)$	A. Weber <i>et al.</i>
$A_{fb}(c)$	A. Weber <i>et al.</i>
A_b	A. Weber <i>et al.</i>
A_c	A. Weber <i>et al.</i>
$A_l(\text{SLD})$	A. Weber <i>et al.</i>
$\sin^2\theta_{\text{eff}}$	A. Weber <i>et al.</i>
m_W	A. Weber <i>et al.</i>
Γ_Z	A. Weber <i>et al.</i>

Improved Heavy flavour code
(particularly for high $\tan\beta$)

Non-Universal Higgs Model (NUHM)



“*NUHM* = *CMSSM* but with decoupled Higgs sector at GUT scale”

 $m_0 := m_{SQ} = m_{SL} \neq m_H$

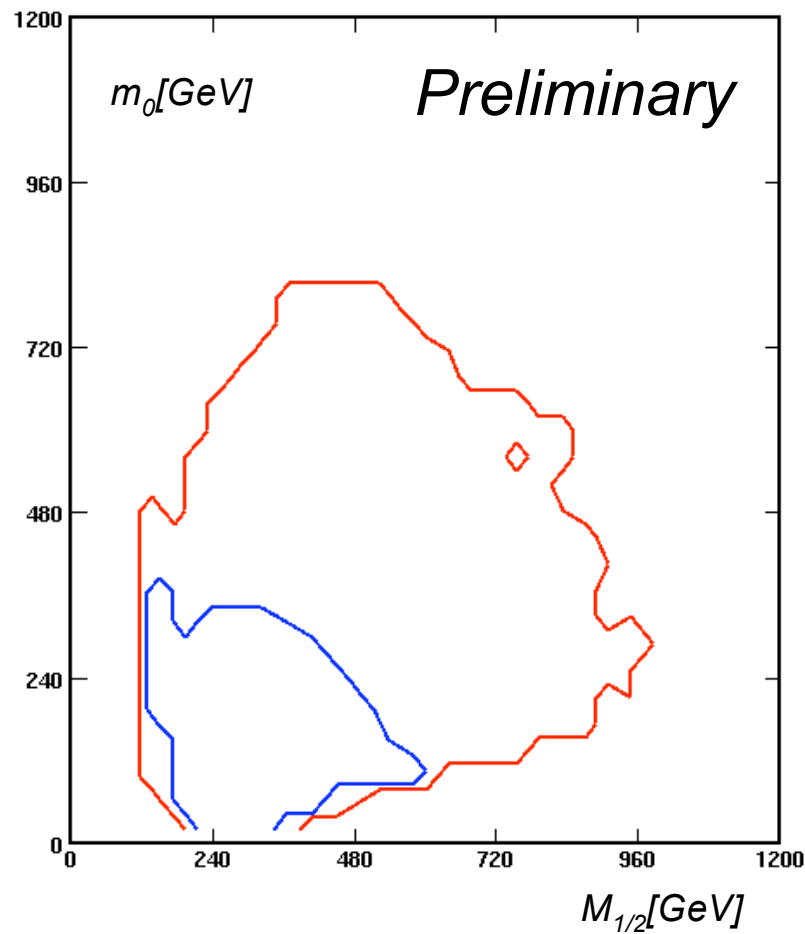
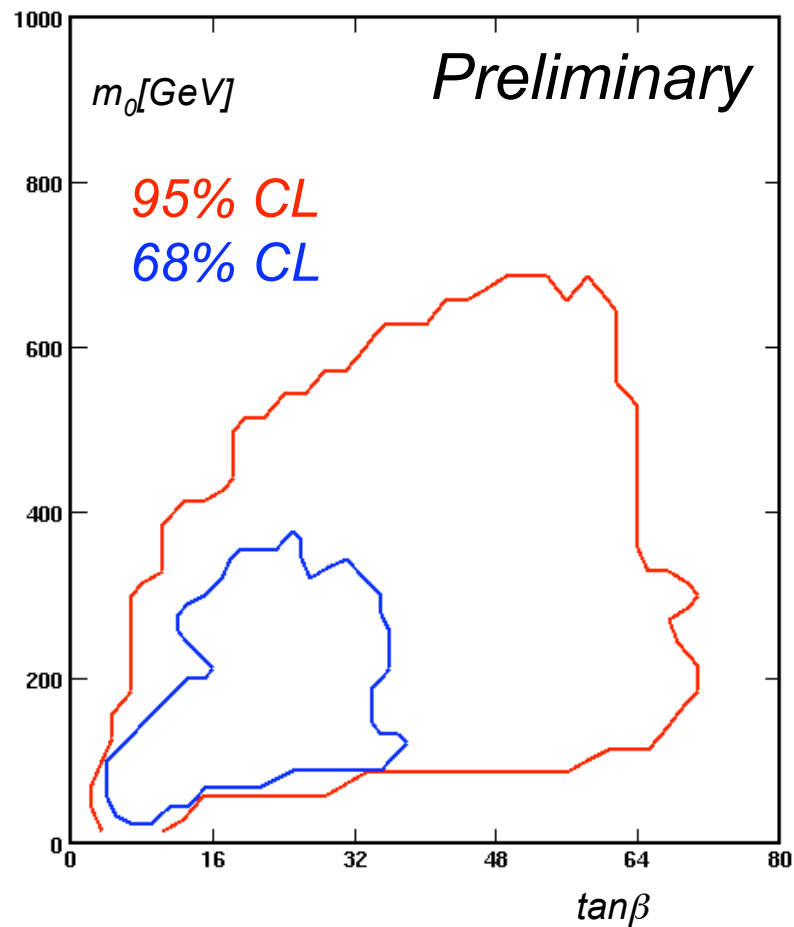
NUHM I Model (5 Parameter)

$m_0, m_{1/2}, A_0, \tan\beta, M_h$ [$M_{hu} = M_{hd}$]

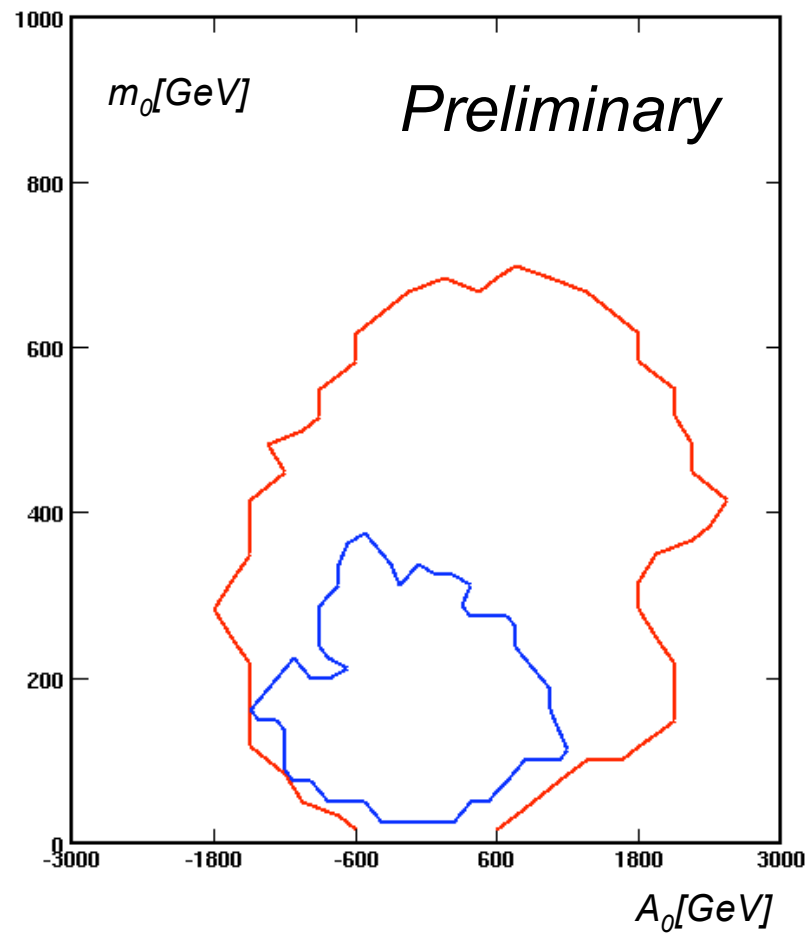
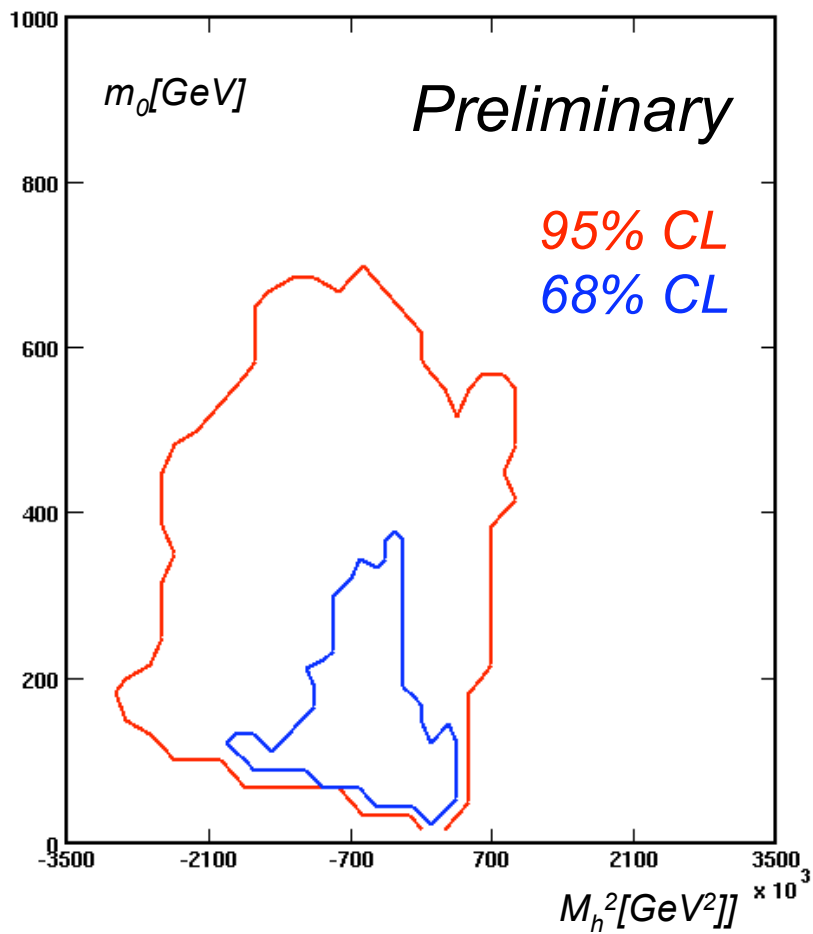
NUHM II Model (6 Parameter)

$m_0, m_{1/2}, A_0, \tan\beta, M_{hu}, M_{hd}$

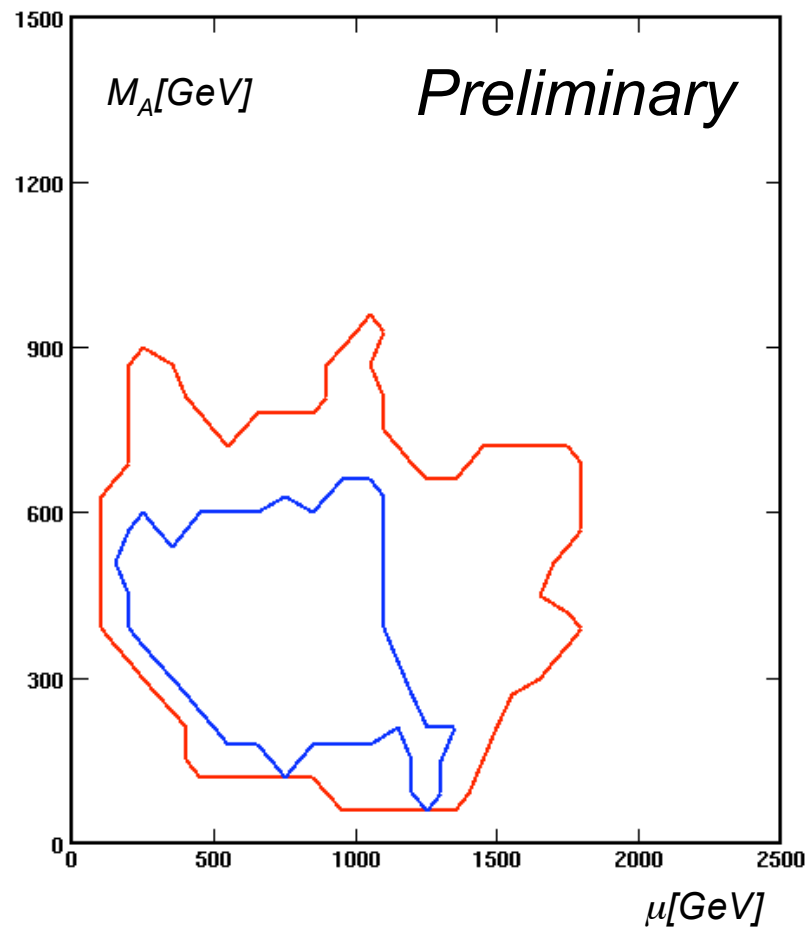
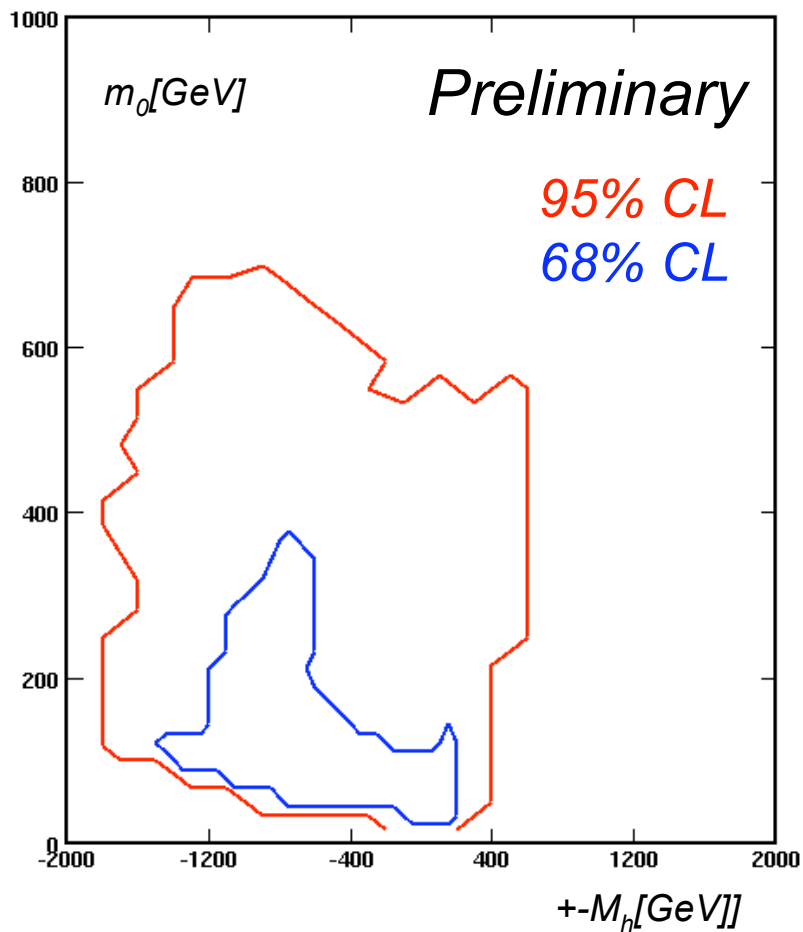
NUHM I Contours - Part I



NUHM I Contours - Part II



NUHM I Contours - Part III



Considered Models



NUMH / TODAY: Use of indirect constraints only

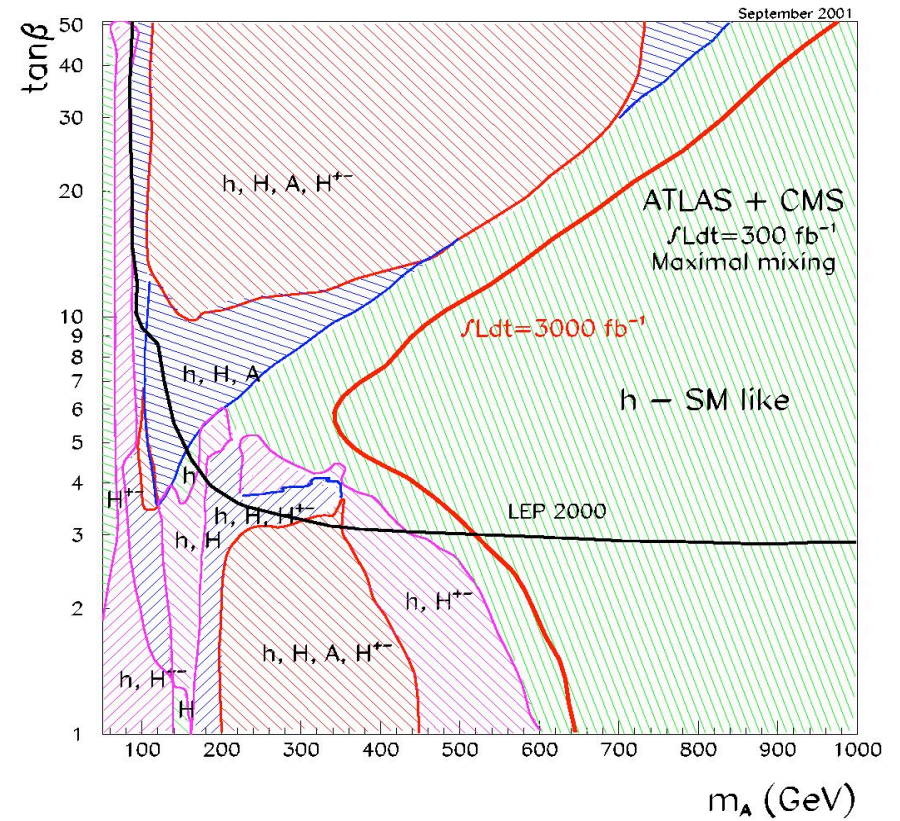
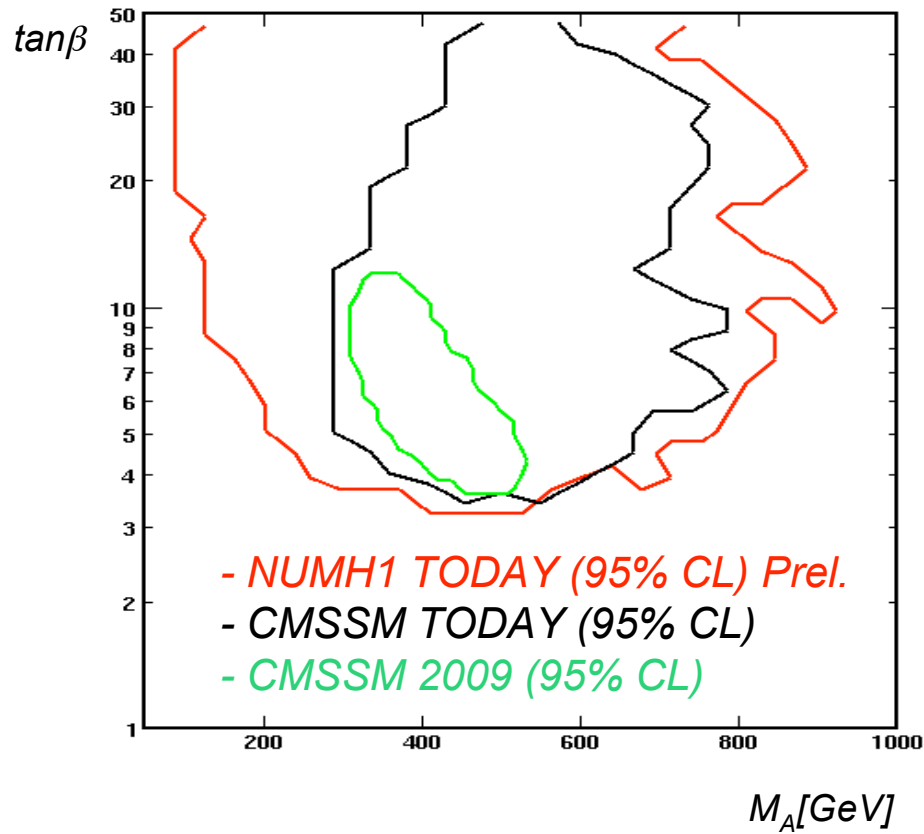
CMSSM TODAY: Use of indirect constraints only

CMSSM 2009: Use of indirect constraints & *assumed kinematic edge measurements from LHC:*

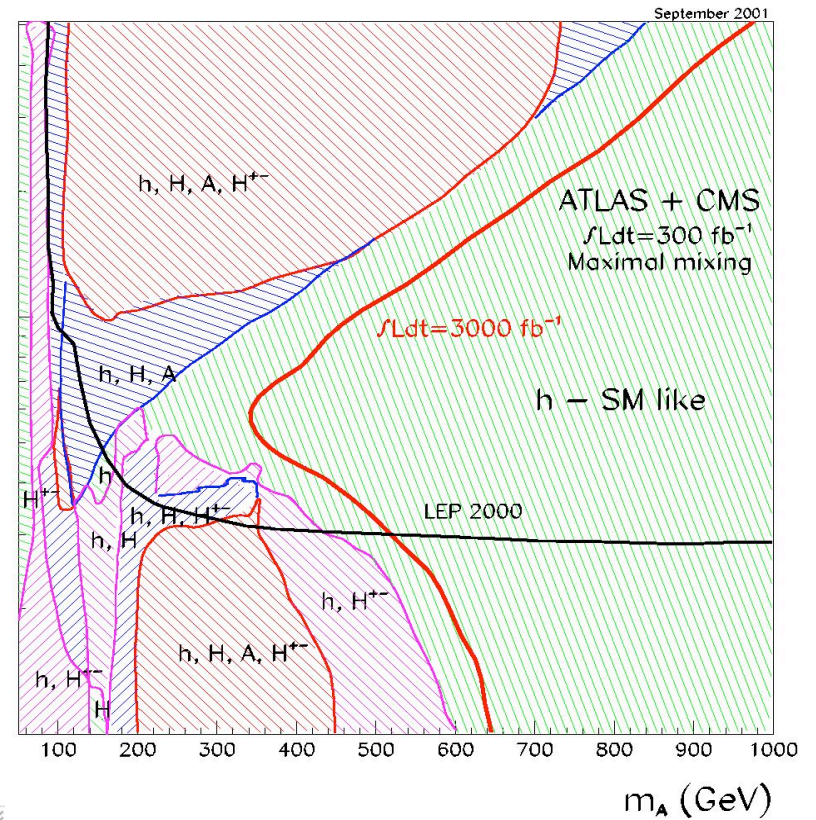
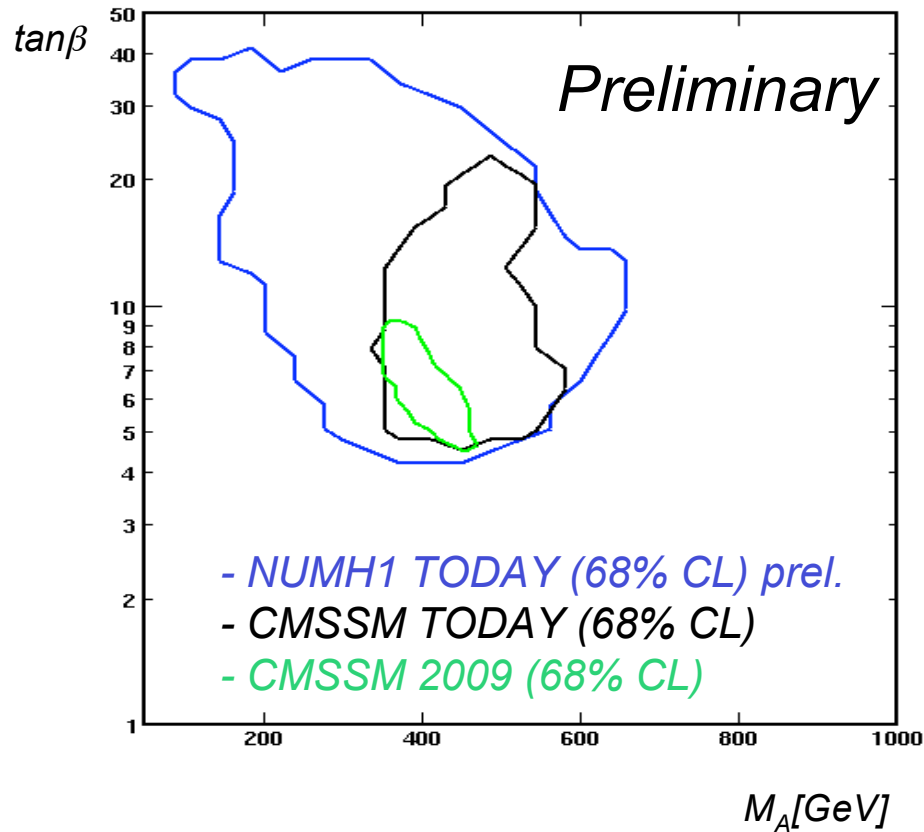
$$\begin{aligned}
 (m_{ll}^2)^{\text{edge}} &= \frac{(m_{\tilde{\chi}_2^0}^2 - m_{\tilde{l}_R}^2)(m_{\tilde{l}_R}^2 - m_{\tilde{\chi}_1^0}^2)}{m_{\tilde{l}_R}^2} \\
 (m_{qll}^2)^{\text{edge}} &= \frac{(m_{\tilde{q}_L}^2 - m_{\tilde{\chi}_2^0}^2)(m_{\tilde{\chi}_2^0}^2 - m_{\tilde{\chi}_1^0}^2)}{m_{\tilde{\chi}_2^0}^2} \\
 (m_{q\tilde{l}}^2)^{\text{edge}}_{\text{min}} &= \frac{(m_{\tilde{q}_L}^2 - m_{\tilde{\chi}_2^0}^2)(m_{\tilde{\chi}_2^0}^2 - m_{\tilde{l}_R}^2)}{m_{\tilde{\chi}_2^0}^2} \\
 (m_{q\tilde{l}}^2)^{\text{edge}}_{\text{max}} &= \frac{(m_{\tilde{q}_L}^2 - m_{\tilde{\chi}_2^0}^2)(m_{\tilde{l}_R}^2 - m_{\tilde{\chi}_1^0}^2)}{m_{\tilde{l}_R}^2}
 \end{aligned}$$

Assume 5% measurement of the edge with leptons only
 Assume 10% measurements of the kinematic Quantities involving jets
 [Conservative uncertainty estimates]

MSSM Higgs -95% CL



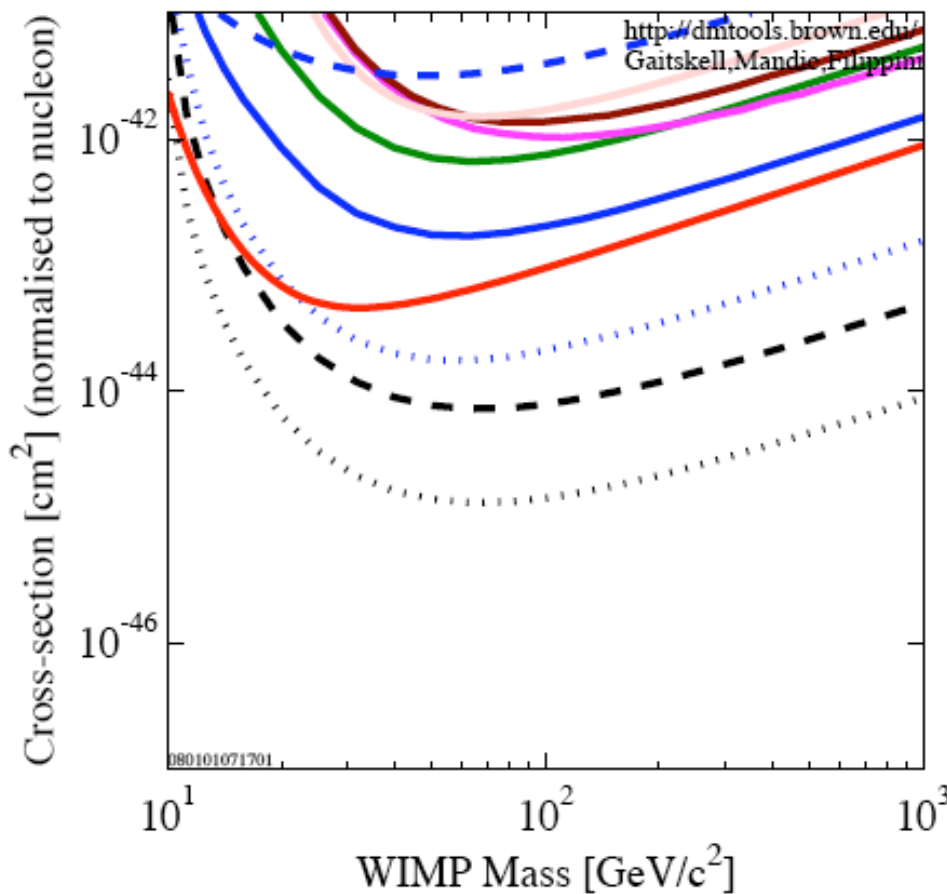
MSSM Higgs - 68% CL



Dark Matter



Direct detection of WIMP (LSP) Dark Matter



- DATA listed top to bottom on plot
- CDMS (Soudan) 2005 Si (7 keV threshold)
 - CRESST 2004 10.7 kg-day CaWO4
 - Edelweiss I final limit, 62 kg-days Ge 2000+2002+2003
 - WARP 2.3L, 96.5 kg-days 55 keV threshold
 - ZEPLIN II (Jan 2007) result
 - CDMS (Soudan) 2004 + 2005 Ge (7 keV threshold)
 - XENON10 2007 (Net 136 kg-d)
 - CDMS Soudan 2007 projected
 - SuperCDMS (Projected) 2-ST@Soudan
 - SuperCDMS (Projected) 25kg (7-ST@Snolab)
- 080101065700

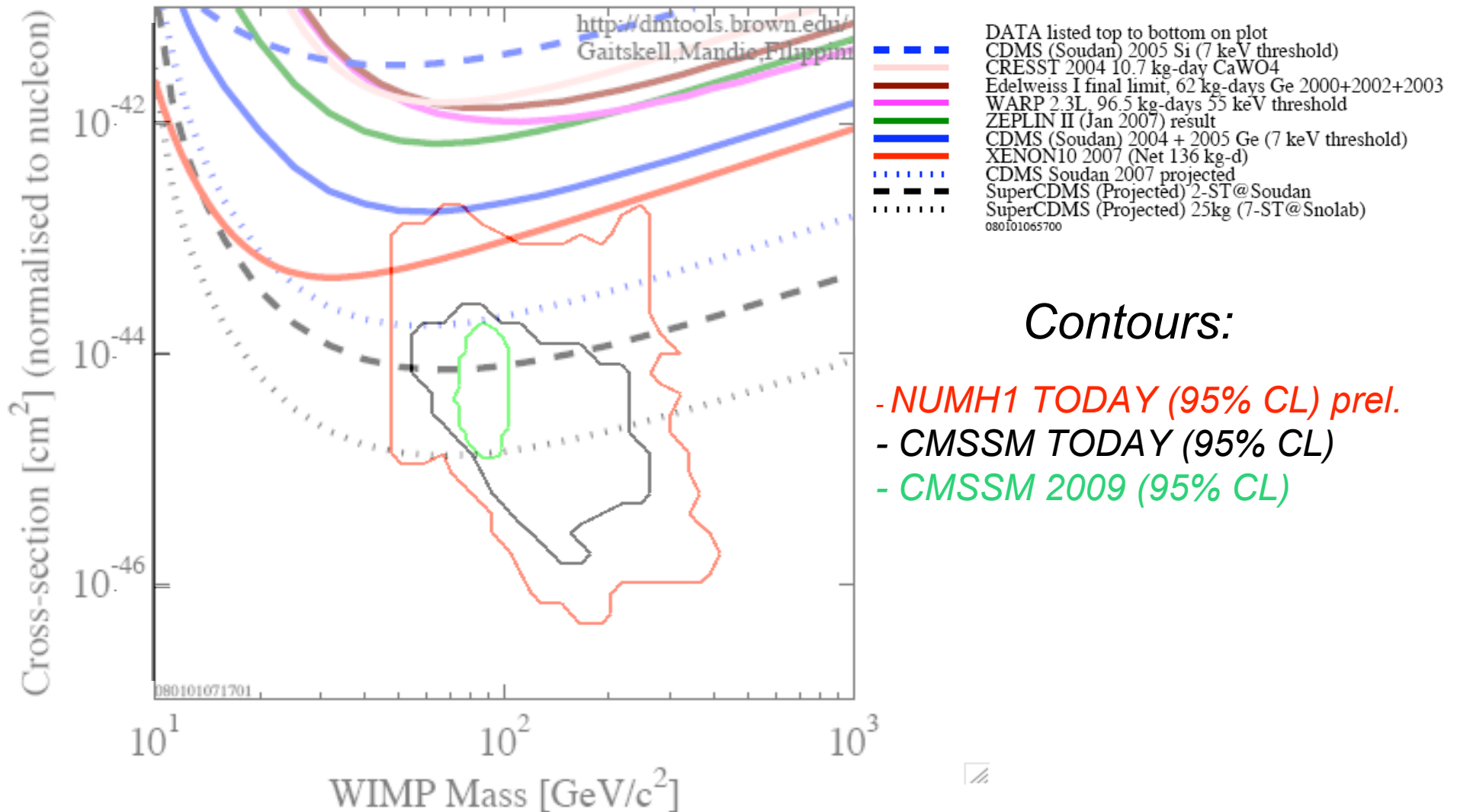
Sensitivity Plot:

WIMP(LSP) Mass vs. σ_p^{SI}

σ_p^{SI} : spin-independent dark matter WIMP elastic scattering cross section on a free proton.

A convenient way to illustrate direct and indirect WIMP searches.

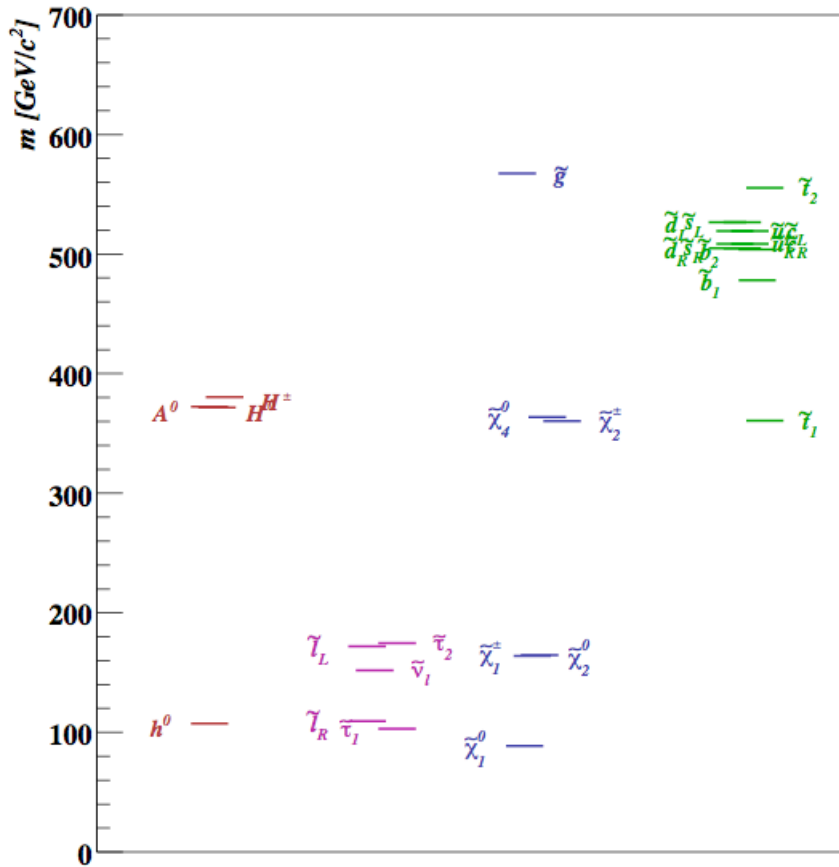
WIMP Sensitivity Plot



Spectrum Comparison

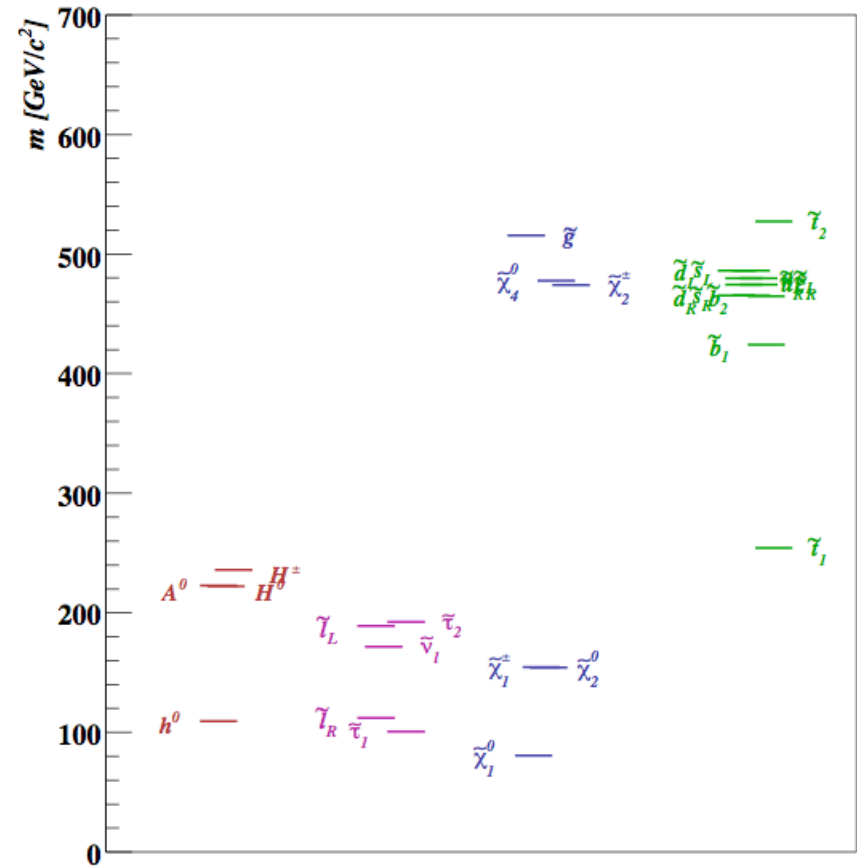


“best CMSSM Fit”



M0 M12 A0 tb
 49.2 232.3 -122.4 6.9
 MA=372 GeV; mu=336 GeV; mh=111 GeV

“best NUHM Fit”



M0 M12 A0 tb Mhd² Mhu²
 101.9 208.1 -523.1 6.7 -183000 -72300
 MA=220 GeV; mu=460 GeV; mh=113 GeV

Conclusions

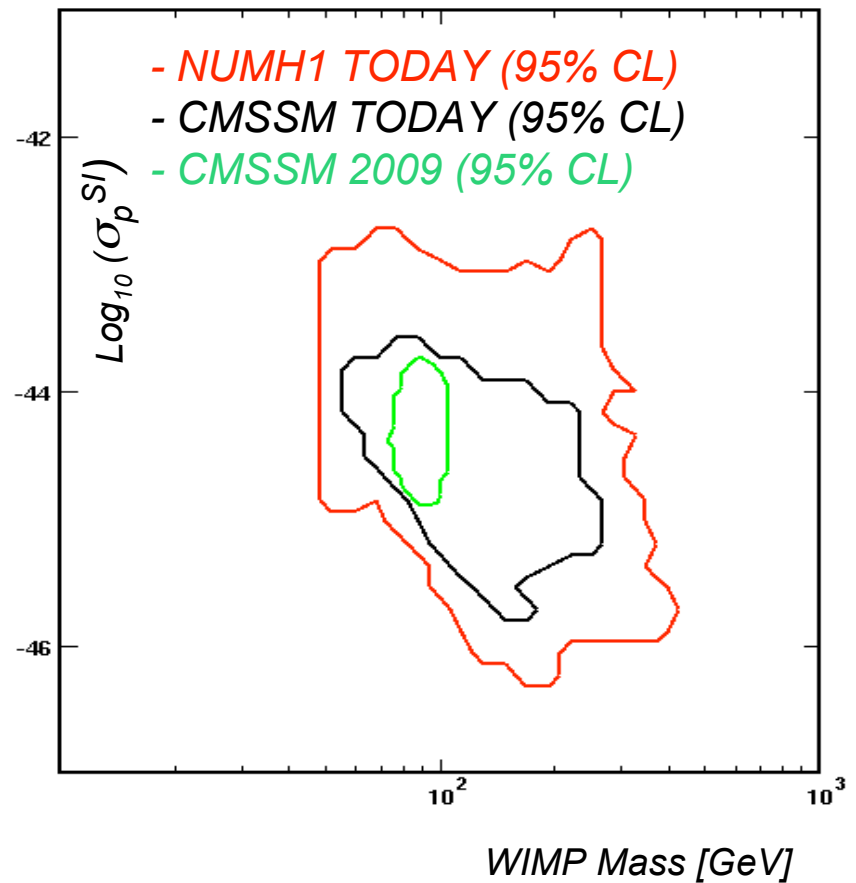
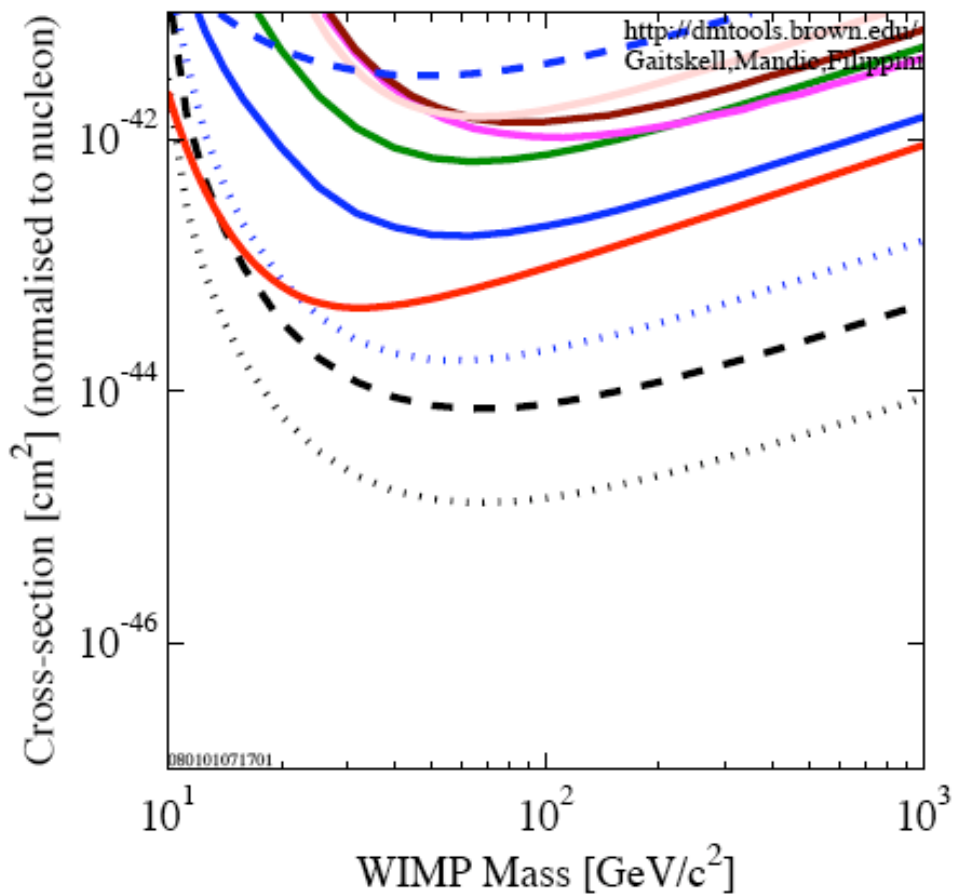


- . Modular framework for comparison (~) in place
 - ✓ Allows to study “preference” of new physics phase space
 - ✓ Allows to study consistency between new signals and precision data
 - ✓ Expect to become important for the interpretation of potential discoveries, eg dark matter and heavy Higgs constraints
- . Early SUSY discovery@LHC “preferred”
- . Higgs perhaps already seen by LEP ☺
- . **This is an open project. Collaborators welcome**
- . Next steps could/will include
 - ✓ More systematic study of the individual effect of different variables
 - ✓ More systematic study on the uncertainties (eg. sparticle spectrum)
 - ✓ More general SUSY models
 - ✓ Other than SUSY BSM



Backup

Dark Matter -WIMP(LSP) contours



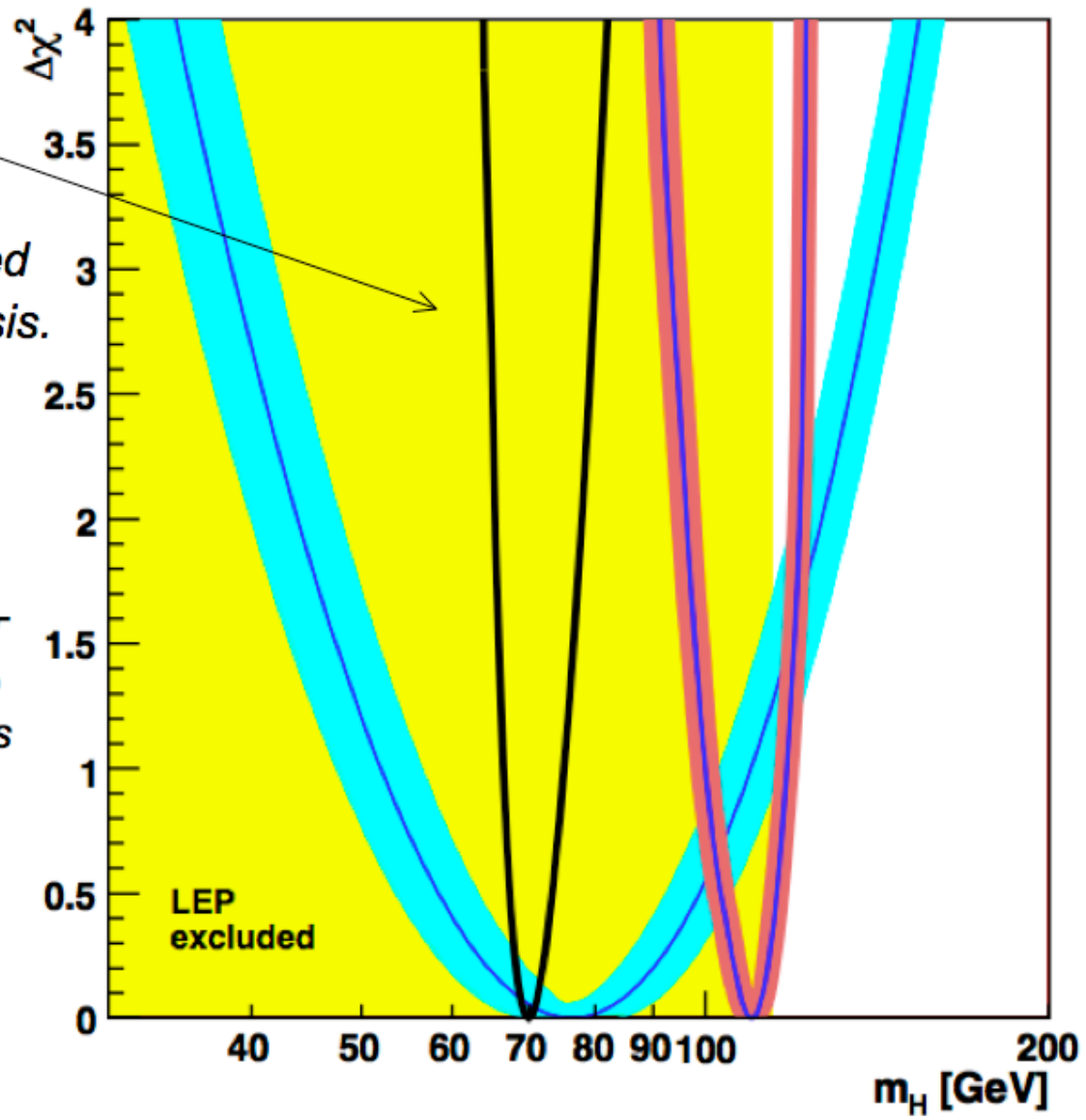
SM-Like Higgs - Sensitivity



Define “hypothetical scenario” with $m_h = 70$ GeV (black line). Errors on indirect constraints are kept but constraints are varied to be compatible with hypothesis.

Conclusion:

“narrowness” of the ellipse is NOT a property of a particular scenario but rather determined by the errors of the indirect constraints and the general model properties of the Higgs sector in the CMSSM.



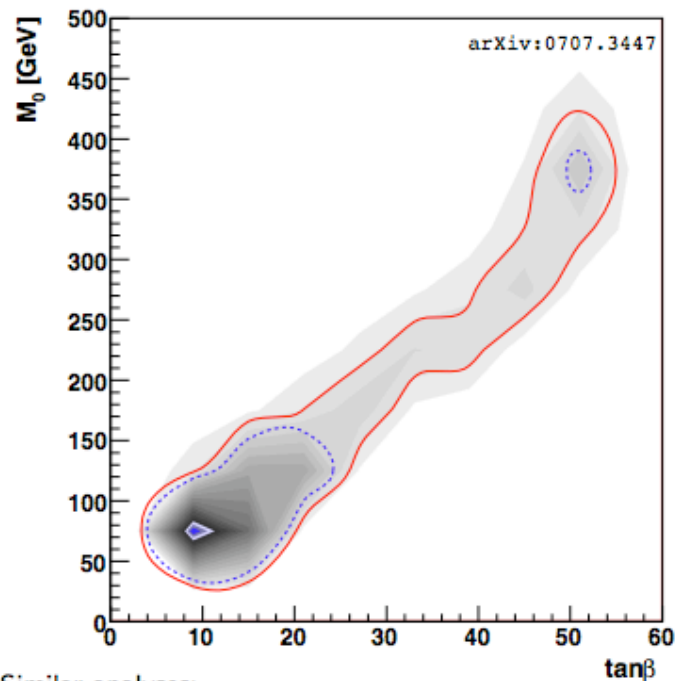
- Multi-parameter χ^2 fit:

$$\chi^2 = \sum_i^N \frac{(C_i - P_i)^2}{\sigma(C_i)^2 + \sigma(P_i)^2} + \sum_i^M \frac{(f_{\text{SM}_i}^{\text{obs}} - f_{\text{SM}_i}^{\text{fit}})^2}{\sigma(f_{\text{SM}_i})^2}$$

- fitting for all CMSSM parameters: M_0 , $M_{1/2}$, A_0 , $\tan \beta$;
 - including relevant SM uncertainties (e.g. m_{top});
- details in O. Buchmüller *et al.*, arXiv:0707.3447 [hep-ph]

Natural extension of J. Ellis *et al.*, arXiv:0706.0652 [hep-ph]

From fits on 2000 pseudo-experiments



Similar analyses:

- Ellis, Heinemeyer, Olive, Weber, Weiglein – hep-ph/0706.0652
- Allanach, Cranmer, Lester, Weber – hep-ph/0705.0487
- Trotta, Austri, Roszkowski – hep-ph/0609126

- overall preferred minimum at low $\tan \beta$, low squark mass;
- less preferred region at high $\tan \beta$, higher squark mass;
- consistent with previous studies.

Note: includes limit from LEP

⇒ Turn to fit *without* limit on m_h
 assessing preferred m_h value
 in CMSSM

Constrain soft-breaking parameters at the GUT scale

CMSSM - a very constraint model:

- Unification of the gaugino [bino, wino and gluino] masses:

$$M_1(M_{\text{GUT}}) = M_2(M_{\text{GUT}}) = M_3(M_{\text{GUT}}) \equiv m_{1/2}$$

- Universal scalar [i.e. sfermion and Higgs boson] masses [i is the generation index]:

$$\begin{aligned} M_{\tilde{Q}_i}(M_{\text{GUT}}) &= M_{\tilde{u}_{Ri}}(M_{\text{GUT}}) = M_{\tilde{d}_{Ri}}(M_{\text{GUT}}) = M_{\tilde{L}_i}(M_{\text{GUT}}) = M_{\tilde{t}_{Ri}}(M_{\text{GUT}}) \\ &= M_{H_u}(M_{\text{GUT}}) = M_{H_d}(M_{\text{GUT}}) \equiv m_0 \end{aligned}$$

- Universal trilinear couplings:

$$A_{ij}^u(M_{\text{GUT}}) = A_{ij}^d(M_{\text{GUT}}) = A_{ij}^l(M_{\text{GUT}}) \equiv A_0 \delta_{ij}$$

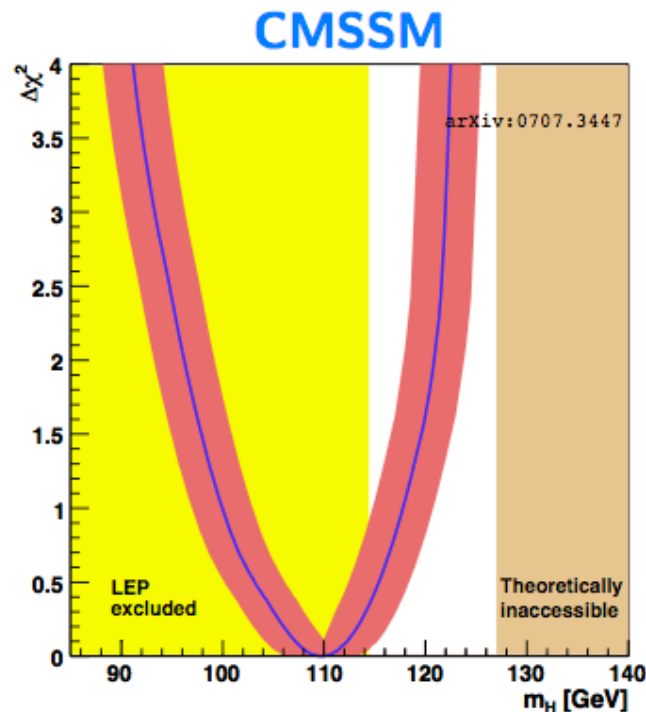
Free Parameters:

$$m_0, m_{1/2}, A_0, \tan\beta, \text{sign}(\mu)$$

All the soft SUSY breaking parameters at the weak scale are obtained through

Renormalization Group Equations (RGE's)

m_0 : common scalar mass at GUT
 $m_{1/2}$: the common gaugino mass at GUT
 $\tan\beta$: $\langle H_u \rangle / \langle H_d \rangle$
 A_0 : common (scalar)³ coupling
 $\text{Sign}(\mu)$: Higgs mass term



- Constrain m_h to scan value;
 - minimize all model parameters in each point;
- ⇒ determine error on m_h prediction

SM fit:

- $m_H = 78_{-24}^{+33} \text{ GeV}/c^2$
- 12% probability at exclusion limit
Including theoretical uncertainty

CMSSM fit:

- $m_h = 110_{-10}^{+8} \pm 3 \text{ GeV}/c^2$
- 20% probability at exclusion limit
Including theoretical uncertainty

New list of constraints

Flavour code

$R(b \rightarrow s\gamma)$	1.13 ± 0.12
$R(\Delta m_s)$	1.040 ± 0.340
$B_s \rightarrow \mu\mu$ [$\times 10^{-8}$]	< 4.7000
$R(B \rightarrow \tau\nu)$	1.07 ± 0.42
$R(B_s \rightarrow X_s ll)$	0.99 ± 0.32
$R(K \rightarrow \tau\nu)$	0.992 ± 0.017
$R(\Delta m_K)$	0.880 ± 0.320
$R(K \rightarrow \pi\nu\nu)$	< 4.5
$B(B_d \rightarrow ll)$ [$\times 10^{-8}$]	< 2.30
$R(\Delta m_d/\Delta m_d)$	1.00 ± 0.08

<i>FeynHiggs</i>	Δa_μ [$\times 10^{-7}$]	2.95 ± 0.87
<i>MicrOMEGAs</i>	Ωh^2	0.113 ± 0.009

SUSY-POPE

$\Delta\alpha_{\text{had}}^{(5)}(m_Z)$	0.02758 ± 0.00035
m_Z [GeV]	91.1875 ± 0.0021
Γ_Z [GeV]	2.4952 ± 0.0023
σ_{had}^0 [nb]	41.540 ± 0.037
R_1	20.767 ± 0.025
A_{fb}^{0j}	0.01714 ± 0.00095
$A_1(P_\tau)$	0.1465 ± 0.0032
R_b	0.21629 ± 0.00066
R_c	0.172 ± 0.003
$A_{\text{fb}}^{0,b}$	0.0992 ± 0.0016
$A_{\text{fb}}^{0,c}$	0.0707 ± 0.0035
A_b	0.923 ± 0.020
A_c	0.670 ± 0.027
$A_1(\text{SLD})$	0.1513 ± 0.0021
$\sin^2\theta_{\text{eff}}^{\text{lept}}(Q_{\text{fb}})$	0.2324 ± 0.0012
m_W [GeV]	80.398 ± 0.025
m_t [GeV]	170.9 ± 1.8

+ $m_h > (115 \pm 1.1 \pm 3)$ GeV from *FeynHiggs*